

A computationally efficient model predictive control scheme for space debris rendezvous

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Space Debris Removal - why

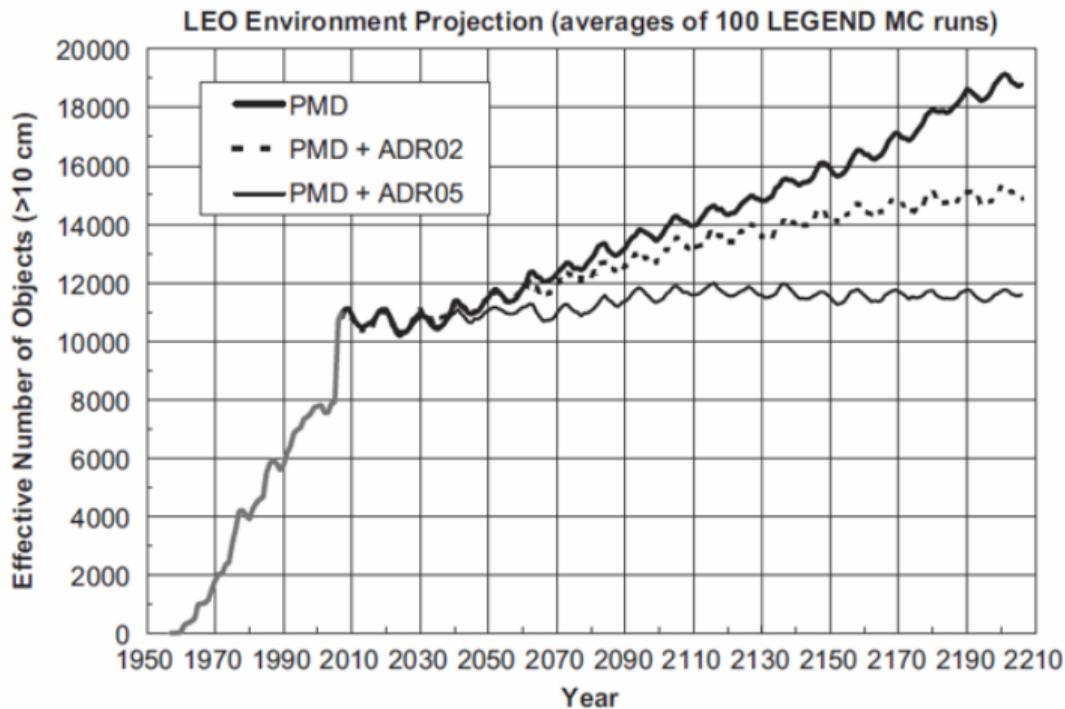
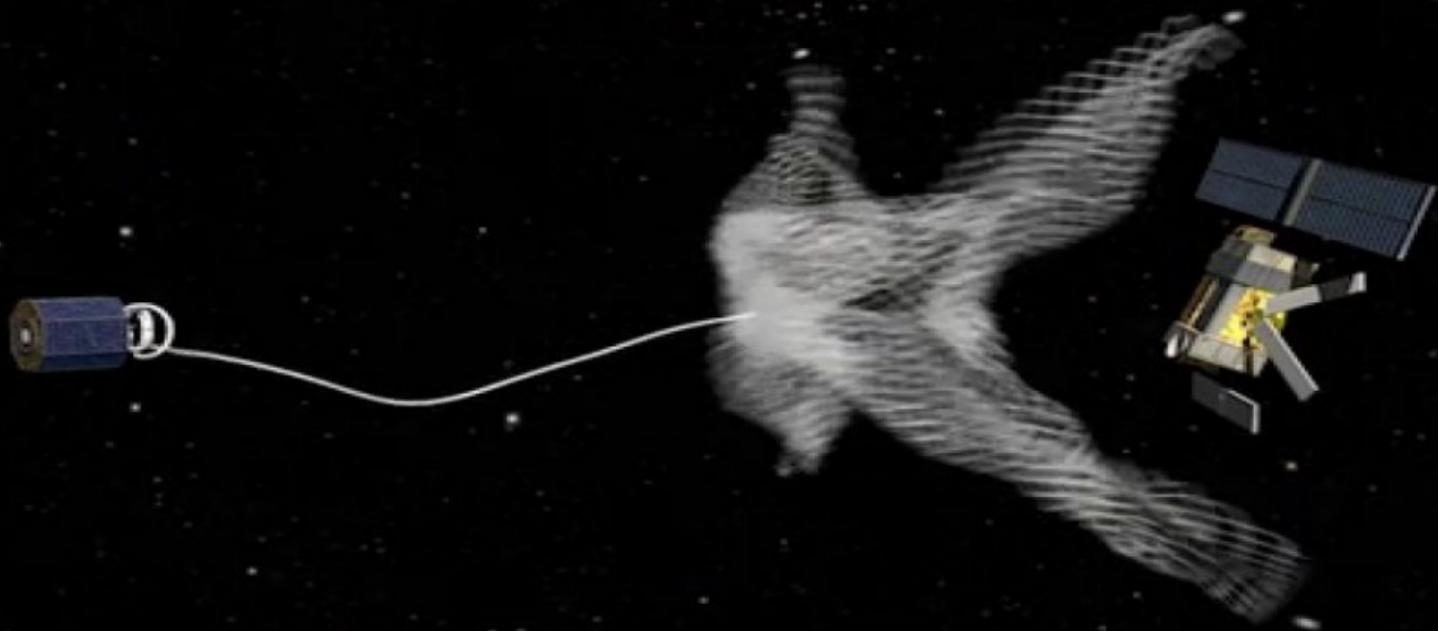


Fig. 2. Comparison of three different scenarios. From top to bottom: postmission disposal (PMD) only, PMD and ADR of two objects per year, and PMD and ADR of five objects per year, respectively.



• esa



overarching need: get spacecrafts close enough to non-collaborative debris

Key challenge for space rendezvous with debris

- better to use small satellites to reduce financial costs
 - ⇒ limited actuation/thrust capabilities
 - ⇒ long maneuvering times

implied need: calculate the on-board thrusters scheduling
taking into account the actuation limits

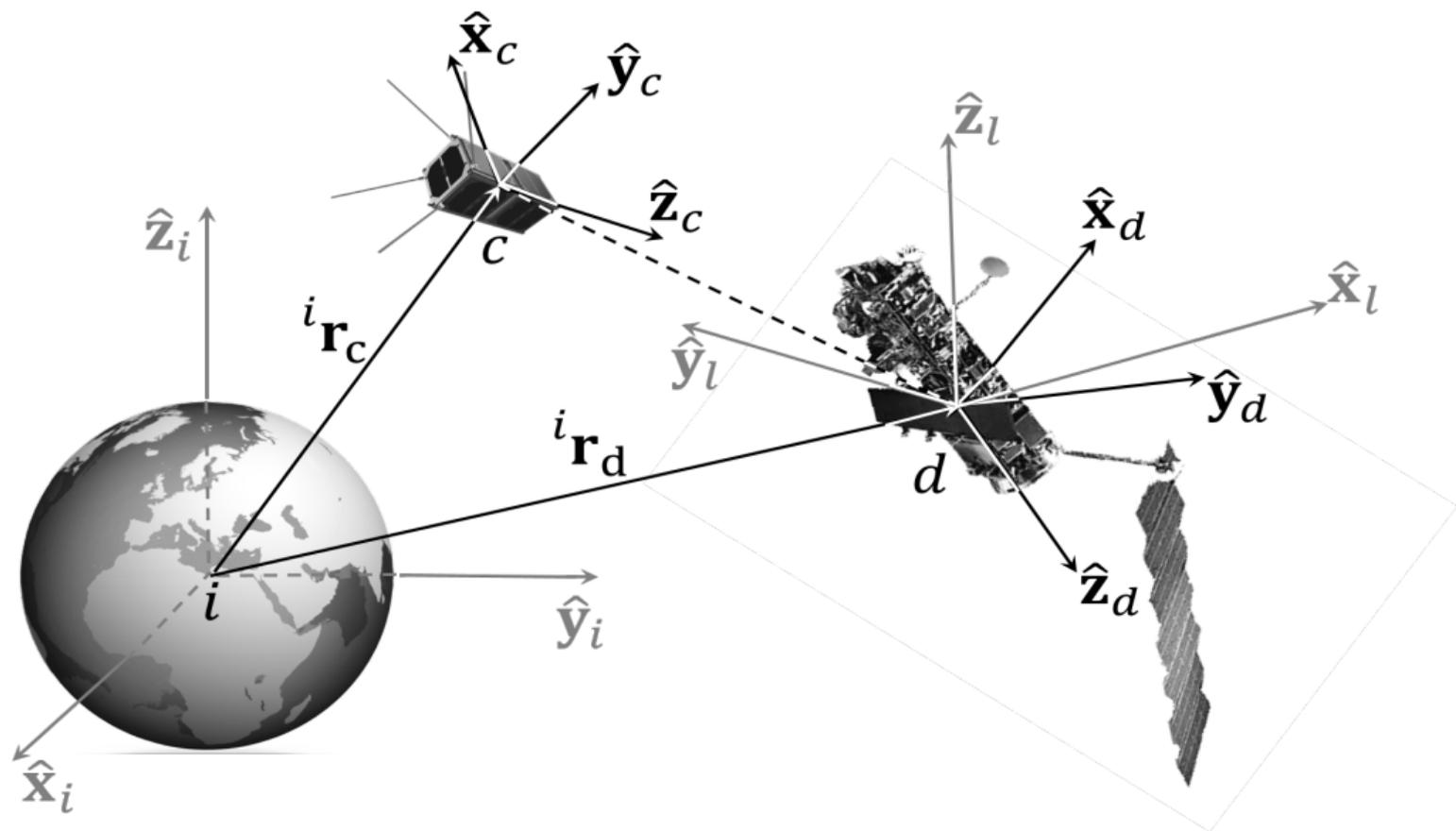
This study (in general)

cast the control problem as a *Model Predictive Control (MPC)* one,
but take into special account the *limited thrusting capabilities* of small satellites

This study (more specifically)

- formulate the MPC so that it:
 - accounts for non-linear orbital dynamics
 - emphasizes fuel consumption minimization
 - is numerically fast, so to achieve longer prediction horizons

Scenario



System Modeling: Orbital Dynamics

- Kinematics: ${}^i \dot{\mathbf{r}}_j = {}^i \mathbf{v}_j$
- Dynamics:

$${}^i \dot{\mathbf{v}}_j = {}^i \mathbf{a}_{g,j} + {}^i \mathbf{a}_{J2,j} + {}^i \mathbf{a}_{drag,j} + {}^i \mathbf{a}_{srp,j} + \frac{{}^i \mathbf{T}_j}{m_j}$$

where:

- $j = \{c, d\}$ (chaser or debris)
- ${}^i \mathbf{a}_{g,j} = -\mu \frac{\mathbf{r}_j}{r_j^3}$ with $\mu =$ Earth's gravitational parameter
- ${}^i \mathbf{a}_{J2,j}$: Earth's oblateness perturbation
- ${}^i \mathbf{a}_{drag,j}$: atmospheric drag perturbation
- ${}^i \mathbf{a}_{srp,j}$: solar radiation pressure perturbation
- m_j : mass
- ${}^c \mathbf{T}_c$: thrust (note: ${}^i \mathbf{T}_d = 0$)

Notation:

- ${}^i \mathbf{r}_j =$ orbital position
- ${}^i \mathbf{v}_j =$ orbital velocity
- $\mathbf{q}_j =$ attitude quaternion
- ${}^j \boldsymbol{\omega}_j =$ angular rate

System Modeling: Attitude Dynamics

- Kinematics: $\dot{\mathbf{q}}_j = \frac{1}{2} {}^j\boldsymbol{\omega}_j \otimes \mathbf{q}_j$
- Dynamics:

$${}^j\dot{\boldsymbol{\omega}}_j = \mathbf{I}_j^{-1}(-{}^j\boldsymbol{\omega}_j \times \mathbf{I}_j {}^j\boldsymbol{\omega}_j + {}^j\boldsymbol{\tau}_{gg,j} + {}^j\boldsymbol{\tau}_j)$$

where:

- \otimes : quaternion product
- \mathbf{I}_j : inertia matrix
- ${}^j\boldsymbol{\tau}_{gg,j}$: gravity gradient torque
- ${}^j\boldsymbol{\tau}_j$: control torque

Notation:

- ${}^i\mathbf{r}_j$ = orbital position
- ${}^i\mathbf{v}_j$ = orbital velocity
- \mathbf{q}_j = attitude quaternion
- ${}^j\boldsymbol{\omega}_j$ = angular rate

note: no torques due to drag and solar radiation pressure

Differences between simulation vs. control models

	<i>simulator</i>	<i>controller</i>
atmospheric drag	✓	
solar radiation pressure	✓	
J_2 effects	✓	
imperfect knowledge about mass distributions	✓	
imperfect actuation	✓	
measurement noise	✓	
mass consumption during maneuvers		

State space representation of the chaser

$$\text{state: } \mathbf{x} := \left[{}^i\mathbf{r}_c^T \quad {}^i\mathbf{v}_c^T \quad \mathbf{q}_c^T \quad {}^c\boldsymbol{\omega}_c^T \quad {}^c\mathbf{T}_c^T \quad {}^c\boldsymbol{\tau}_c^T \right]^T$$

$$\text{inputs: } \mathbf{u} := \left[{}^c\dot{\mathbf{T}}_c^T \quad {}^c\dot{\boldsymbol{\tau}}_c^T \right]^T$$

in words:

state = orbital position & velocity,
attitude, angular rate, applied thrusts and torques

inputs = rate of variation of the applied thrusts and torques

Mission = make the chaser match its orbit with that of the debris

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$$\implies \text{relative dynamics: } \begin{cases} {}^l \mathbf{d}_c := {}^l \mathbf{R}_i ({}^i \mathbf{r}_c - {}^i \mathbf{r}_d) & (\text{position}) \\ {}^l \dot{\mathbf{d}}_c := {}^l \mathbf{R}_i ({}^i \mathbf{v}_c - {}^i \mathbf{v}_d) + {}^i \boldsymbol{\omega}_l \times {}^l \mathbf{R}_i ({}^i \mathbf{r}_c - {}^i \mathbf{r}_d) & (\text{velocity}) \\ \mathbf{q}_{err} := \mathbf{q}_{c,ref} \otimes \mathbf{q}_c^* & (\text{attitude error}) \end{cases}$$

with

${}^c \mathbf{R}_i$ = rotation matrix from the ECI to the LHLV reference frame

${}^i \boldsymbol{\omega}_l$ = angular velocity of the LHLV with respect to the ECI reference frame

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controller's goal = make $\mathbf{y} := [{}^l \mathbf{d}_c \quad {}^l \dot{\mathbf{d}}_c \quad \mathbf{q}_{err}] \rightarrow \mathbf{0}$

Formulation as an optimal control problem

$$\min_{\hat{\mathbf{u}}} \int_0^T (\|\hat{\mathbf{y}}\|_Q^2 + \|\hat{\mathbf{u}}\|_R^2) d\tau$$

$$\text{s.t. } \hat{\mathbf{x}}_0 = \mathbf{x}(t)$$

$$\hat{\dot{\mathbf{x}}} = f(\hat{\mathbf{x}}, \hat{\mathbf{u}})$$

$$\hat{\mathbf{y}} = h(\hat{\mathbf{x}}, {}^i\hat{\mathbf{r}}_d, {}^i\hat{\mathbf{v}}_d, \mathbf{q}_{c,ref})$$

$${}^c\hat{\mathbf{T}}_c \in [{}^c\underline{\mathbf{T}}_c \quad {}^c\overline{\mathbf{T}}_c]$$

$${}^c\hat{\boldsymbol{\tau}}_c \in [{}^c\underline{\boldsymbol{\tau}}_c \quad {}^c\overline{\boldsymbol{\tau}}_c]$$

$${}^c\hat{\boldsymbol{\omega}}_c \in [{}^c\underline{\boldsymbol{\omega}}_c \quad {}^c\overline{\boldsymbol{\omega}}_c]$$

$${}^c\hat{\dot{\mathbf{T}}}_c \in [{}^c\underline{\dot{\mathbf{T}}}_c \quad {}^c\overline{\dot{\mathbf{T}}}_c]$$

$${}^c\hat{\dot{\boldsymbol{\tau}}}_c \in [{}^c\underline{\dot{\boldsymbol{\tau}}}_c \quad {}^c\overline{\dot{\boldsymbol{\tau}}}_c]$$

→ computational demands

... in discrete time

$$\min_{x_k, u_k} \sum_{k=0}^{N-1} d(h(x_k, u_k) - h_{ref}^k)_{W_k} + d_N(h_N(x_N) - h_{ref}^N)_{W_N}$$

$$\text{s.t. } 0 = x_0 - \hat{x}_0$$

$$0 = x_{k+1} - \phi_k(x_k, u_k)$$

$$\underline{x}_k \leq x_k \leq \bar{x}_k$$

$$\underline{u}_k \leq u_k \leq \bar{u}_k$$

$$\underline{r}_k \leq r_k(x_k, u_k) \leq \bar{r}_k$$

with $\phi_k(x_k, u_k)$ a numerical integrator simulating

$$0 = f(\dot{x}(t), x(t), u(t), t) \quad x(0) = x_k$$

Our strategy: use MATMPC

Matlab
(relatively slow)



C++ libraries
(faster but less straightforward)

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C++ libraries
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MATMPC:

- based on MATLAB
- modular structure
- core modules written in MATLAB C API
- not a C/C++ library, but a combination of MATLAB scripts and MEX functions
- computationally fast (comparable to low level languages)

Supported algorithms

Hessian Approximation	Generalized Gauss-Newton		
Integrator	Explicit Runge Kutta 4 (CasADi code generation)	Explicit Runge Kutta 4	Implicit Runge-Kutta (Gauss-Legendre)
Condensing	non	full	partial
QP solver	qpOASES OSQP	MATLAB quadprog HPIP	Ipopt
Globalization	ℓ_1 merit function	line search	Real-Time Iteration
Special features	CMoN-SQP	input move blocking	non-uniform grid

implication: can consider longer prediction horizons
⇒ better divide the mission in distance-depending stages

Indeed:

when too far from the debris...

- ...info about the target's motion are of limited quality
⇒ following exactly the reference trajectory may be fuel-inefficient
& better to have longer sampling periods

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chosen approach: use two MPCs with the same dynamics
and different prediction horizons / cost functions

What do we want to understand?

How do...

- the length of the prediction horizon
- the definition of "close" vs "far"

...affect the mission efficiency? (*fuel and time*)

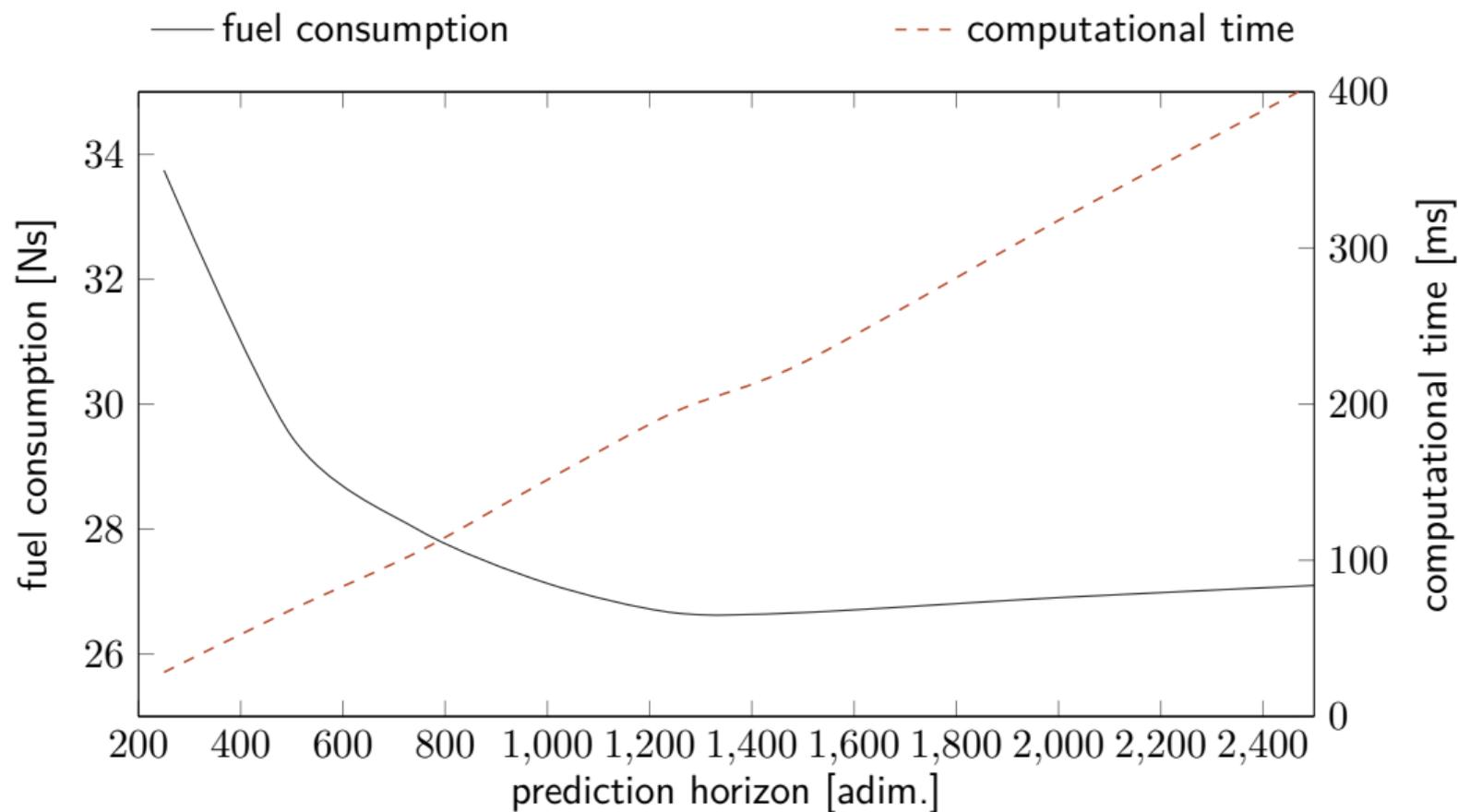
Simulation Parameters - scenario

- Space debris:
 - in circular orbit
 - altitude: $300km$
 - inclination: 30 deg
- Chaser:
 - 3U CubeSat
 - same orbital parameters as space debris
 - mass: $4kg$
 - inertia matrix:

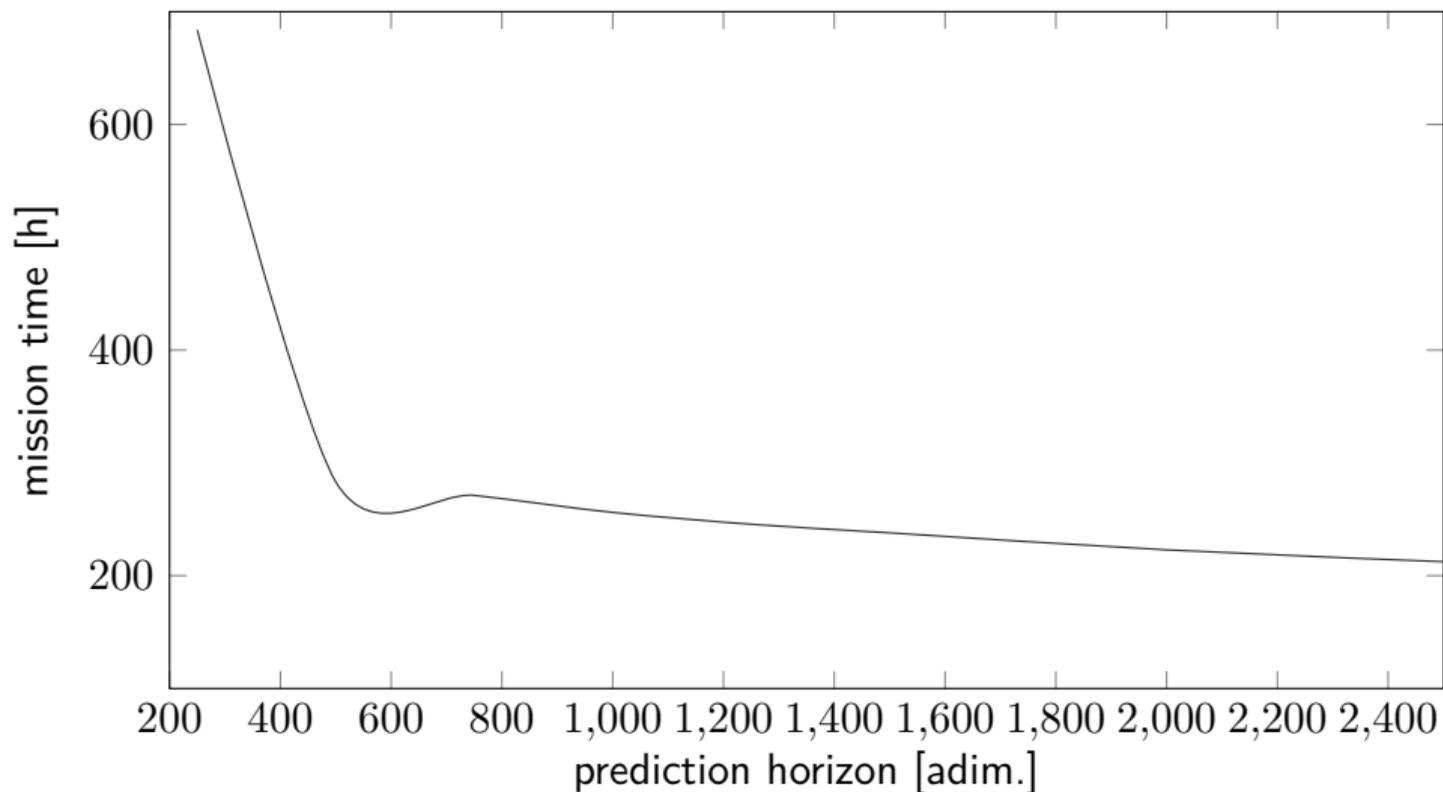
$$I_c = \begin{bmatrix} 0.0333 & 0.0000 & 0.0000 \\ 0.0000 & 0.0067 & 0.0000 \\ 0.0000 & 0.0000 & 0.0333 \end{bmatrix} kg\ m^2$$

- *true anomaly offset*: $\Delta\nu = -10\text{ deg} \implies$ *initial distance from the debris*: $1164km$

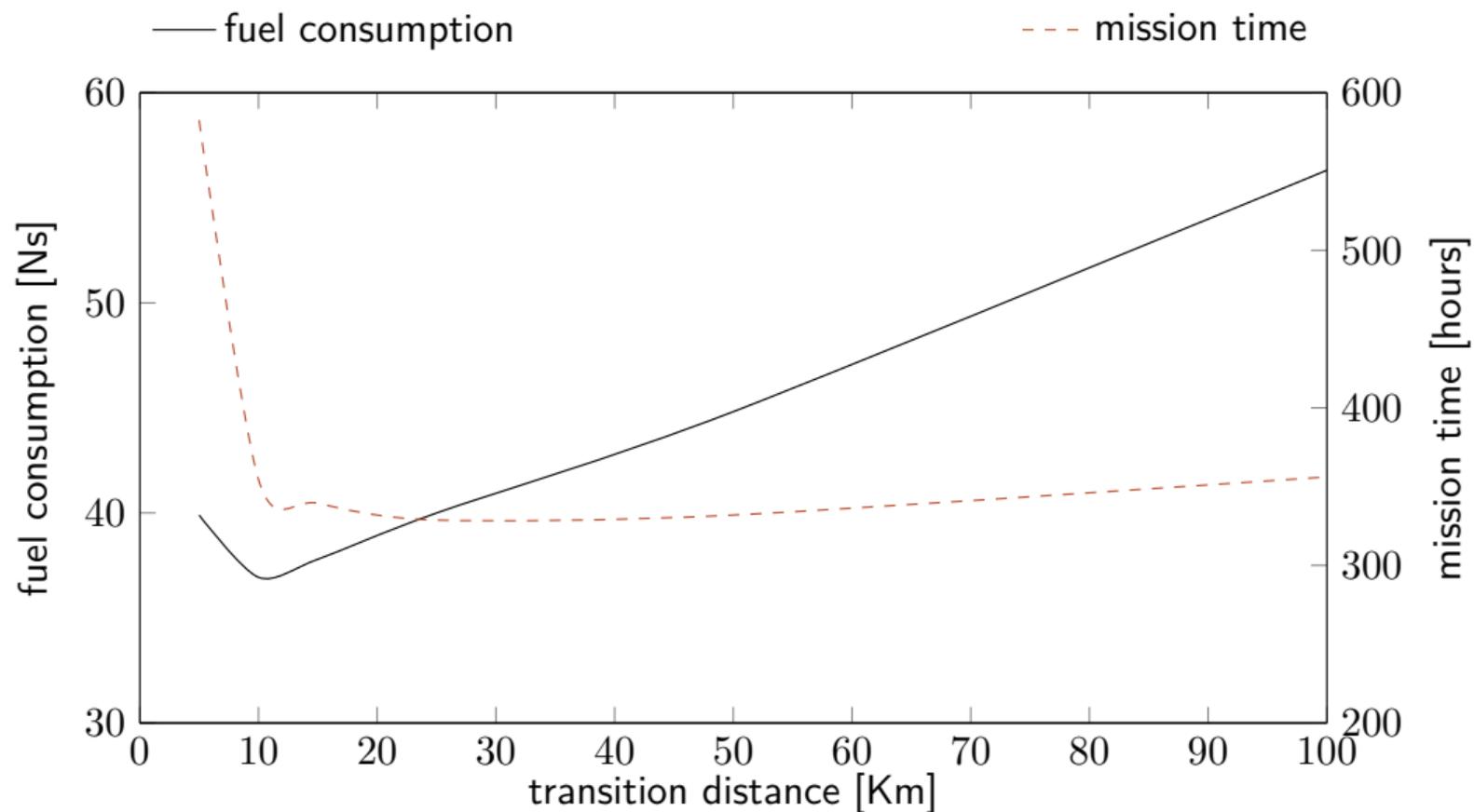
Prediction horizon length vs mission efficiency (1)



Prediction horizon length vs mission efficiency (2)



Transition distance vs mission efficiency



Conclusions

- if having low actuation capabilities, then better to increase the prediction-horizons
- long prediction-horizons require computational efficiency
- when considering long missions we should use time-varying controllers
(\implies *worth to be explored further*)

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